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Steps towards the Well-posedness of the Characteristic Evolution for the Einstein Equations

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Steps towards the Well-posedness of the Characteristic Evolution for the Einstein Equations

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APS April Meeting
Denver, CO, April 16, 2013

arXiv:1106.4841, Phys.Rev.D84:044057,2011

<i>Variable</i>	<i>Re</i>	<i>Im</i>
$\mathcal{E}_2(\Psi)_{R=20}$	1.14×10^{-3}	1.17×10^{-3}
$\mathcal{E}_2(\Psi)_{R=50}$	4.04×10^{-4}	3.53×10^{-4}
$\mathcal{E}_2(\Psi)_{R=100}$	2.81×10^{-4}	2.09×10^{-4}
$\mathcal{E}_2(\delta\psi)_{R=50}$	5.09×10^{-3}	5.08×10^{-3}
$\mathcal{E}_2(\delta\psi)_{R=100}$	6.81×10^{-3}	6.32×10^{-3}
$\mathcal{E}_2(\Psi_{\Delta R(50,100)})$	1.94×10^{-2}	1.91×10^{-2}
$\mathcal{E}_2(\psi_{4,\Delta R(50,100)})$	3.13×10^{-2}	3.14×10^{-2}

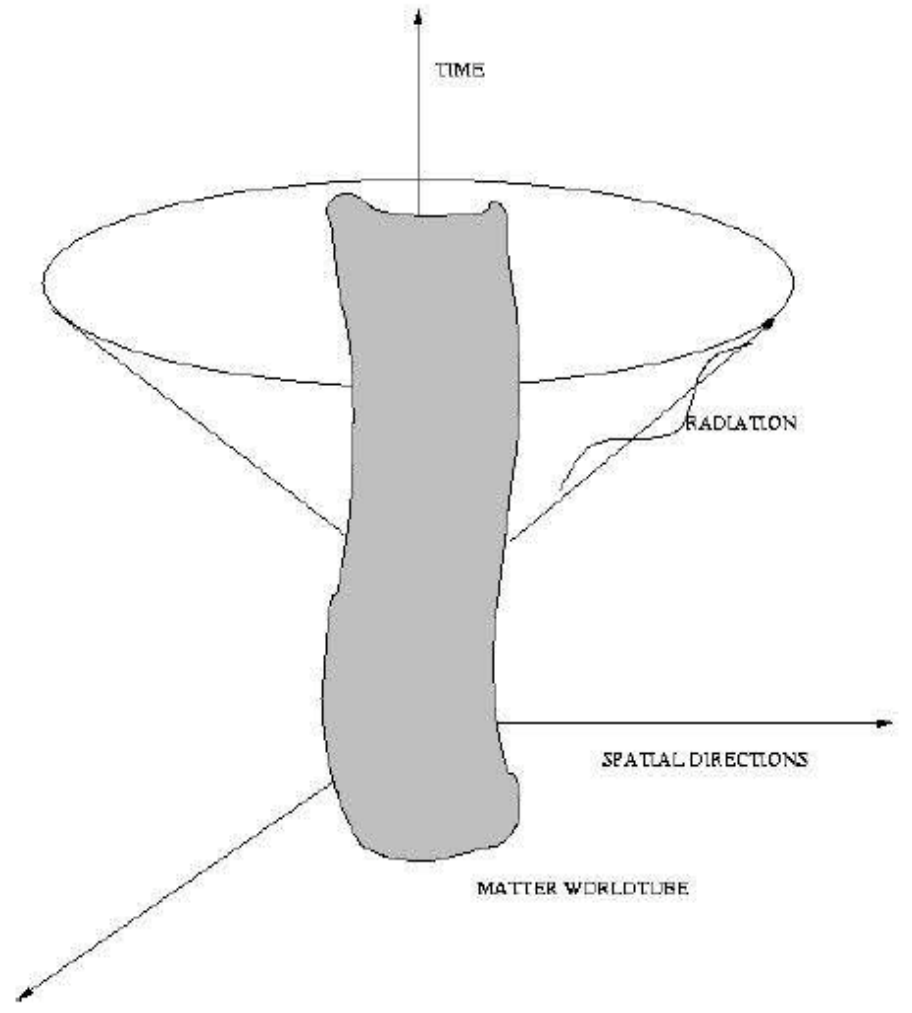
*

The new Characteristic Extraction Module satisfies the criteria required by Advanced LIGO for detection and measurement.

*


Publicly available as part of Einstein Toolkit

Background



- * Overall 1.5 accuracy ought to be improved
- * Independent error sources
- * The only way to ensure both the stability and the accuracy of the simulation is a well-posed algorithm.
- * A new characteristic code would be of great value.

Idea


$$g^{ab} \nabla_a \nabla_b \Phi = S,$$

$$\partial_t \partial_x \Phi = S,$$

$$(x = \tilde{t} + \tilde{x}, t = \tilde{t} - \tilde{x})$$

$$\Phi(x, t) \rightarrow \Phi(x, t) + g(t) *$$

* The well-posedness of the null-timelike problem for the Einstein equations has not yet been established.

* The wave equation in characteristic coordinates allows solution freedom independent of initial data.

Problem



$$\Phi = e^{ax} \Psi \Rightarrow \Psi = e^{-ax} \Phi, a > 0$$

H-O Kreiss and J. Winicour, CQG 28, 2011

$$\partial_t(\partial_x \Phi + a\Phi) = \partial_y^2 \Phi - 2b\partial_y \Phi + S$$

$$\partial_t(\partial_x \Phi + a\Phi) = (\partial_x^2 + \partial_y^2)\Phi - 2b\partial_y \Phi + S$$

$$\partial_x \rightarrow D_{0x}, \partial_y \rightarrow D_{0y},$$

$$\partial_x^2 \rightarrow D_{+x}D_{-x}, \partial_y^2 \rightarrow D_{+y}D_{-y}$$

$$\Phi(t, j_1, j_2) = \frac{1}{N} \sum_0^{N-1} \sum_0^{N-1} \hat{\Phi}(t, f_1, f_2) e^{i2\pi h(j_1 f_1 + j_2 f_2)}$$

* Change of variable, introduce the *a-term*, and bound the growing modes

* Reduce the problem to Cauchy (whole space):

* Double-null

* Null-timelike

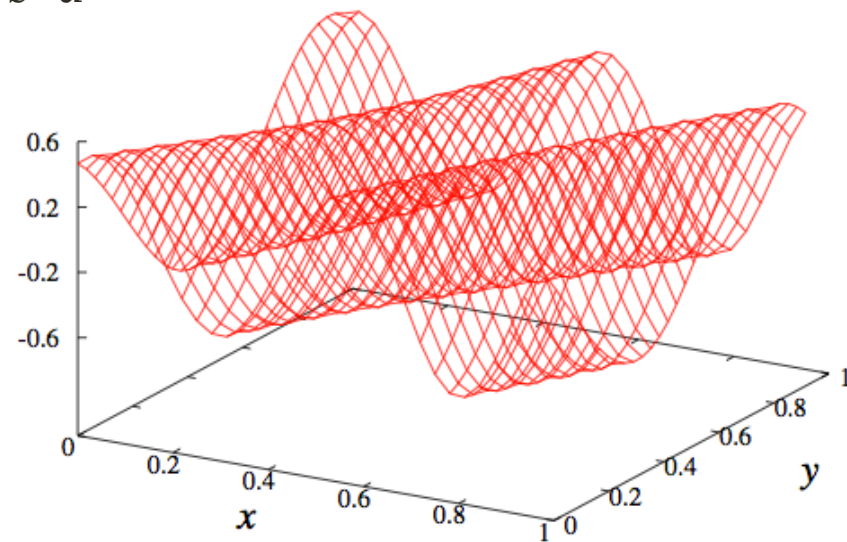
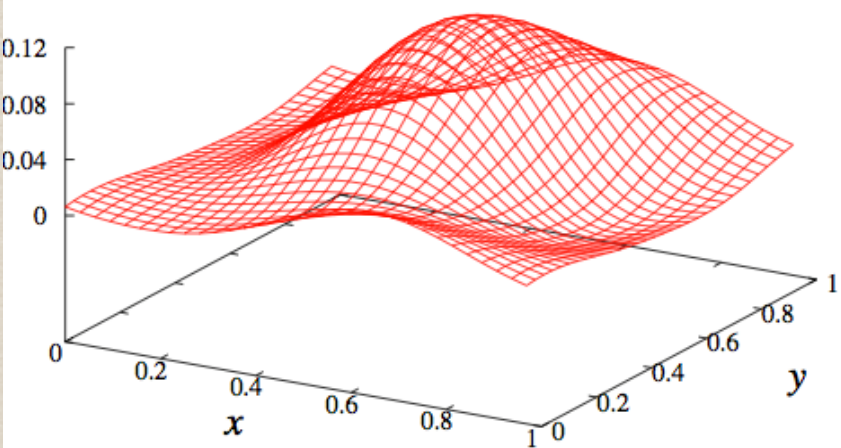
* Analyze stability against lower-order perturbations for the numeric problem

Method

$\Phi(t,x,y)$ at $t=1, b=0.5$

Double-null, $b \approx a$

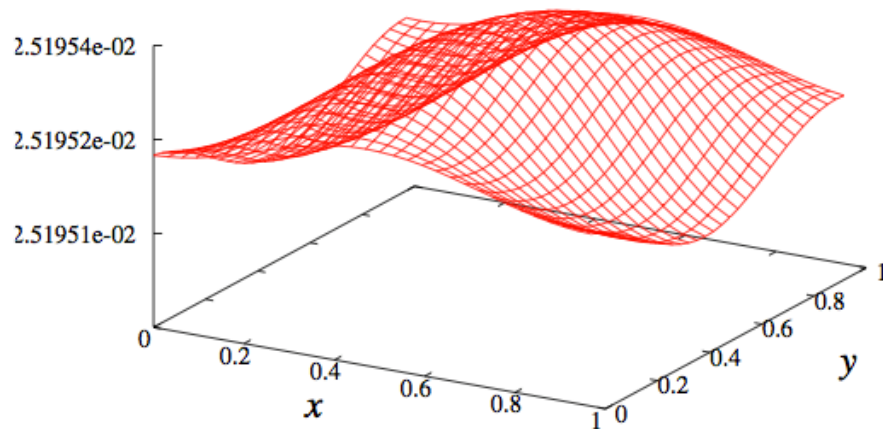
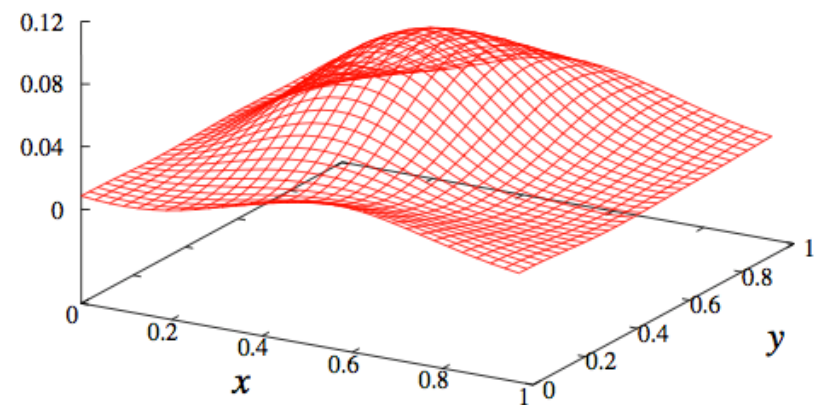
$\Phi(t,x,y)$ at $t=25, b=0.5$



$\Phi(t,x,y)$ at $t=1, b=0.5$

Null-timelike, $b \approx a$

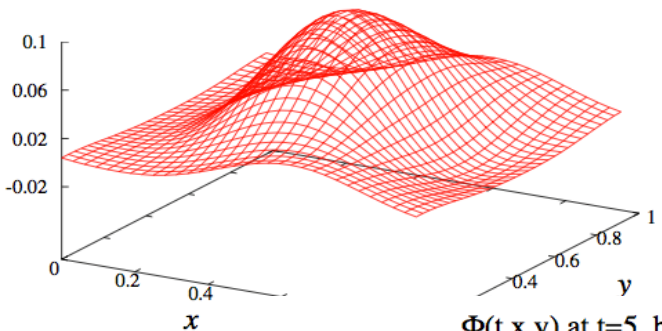
$\Phi(t,x,y)$ at $t=25, b=0.5$



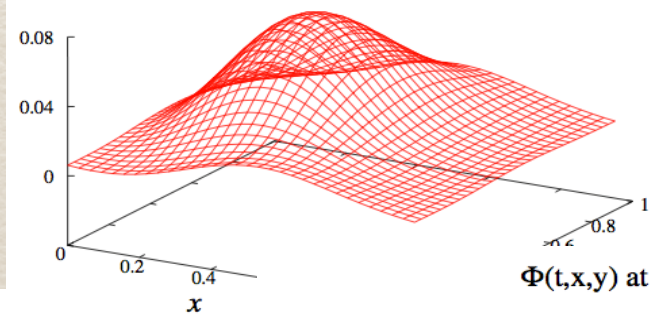
Whole-space



$\Phi(t,x,y)$ at $t=1, b=0.1$

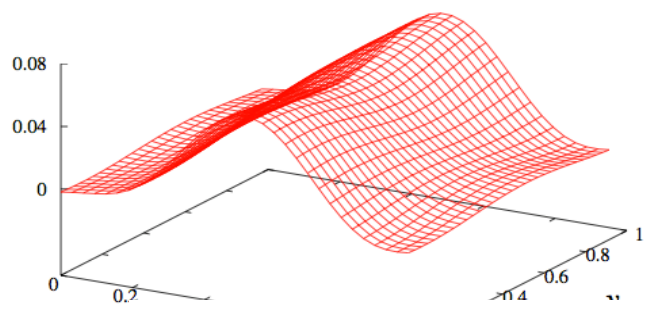


$\Phi(t,x,y)$ at $t=1$

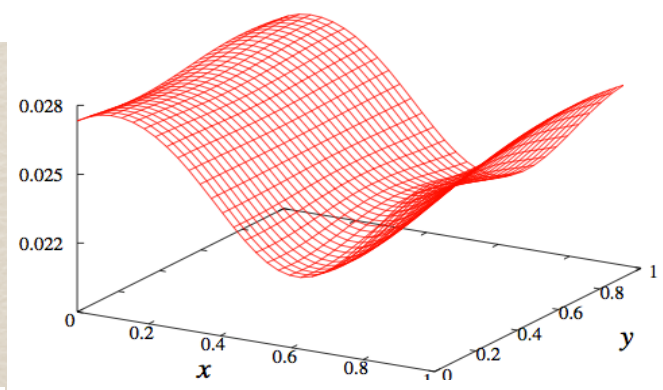


Null-timelike, $b < a$

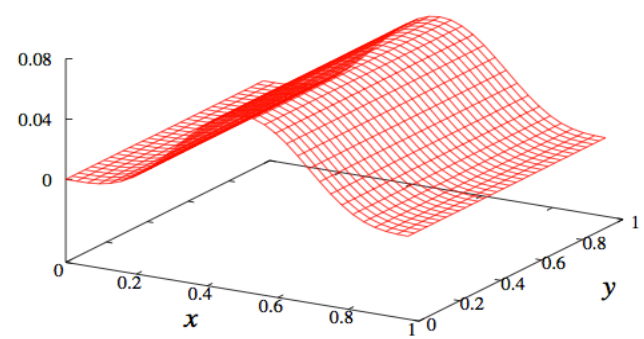
$\Phi(t,x,y)$ at $t=5, b=0.1$



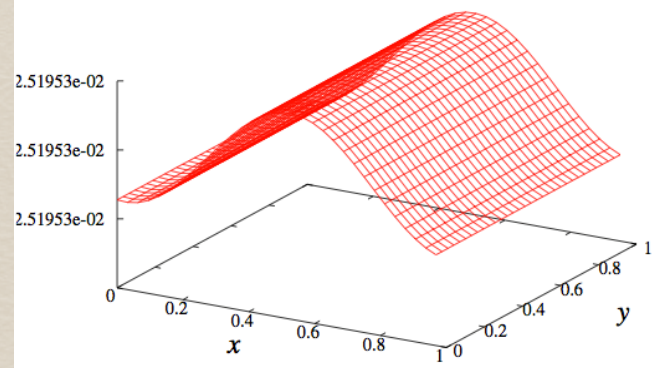
$\Phi(t,x,y)$ at $t=5$



$\Phi(t,x,y)$ at $t=50, b=0.1$



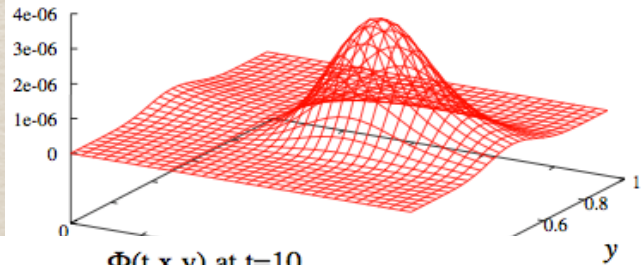
$\Phi(t,x,y)$ at $t=50$



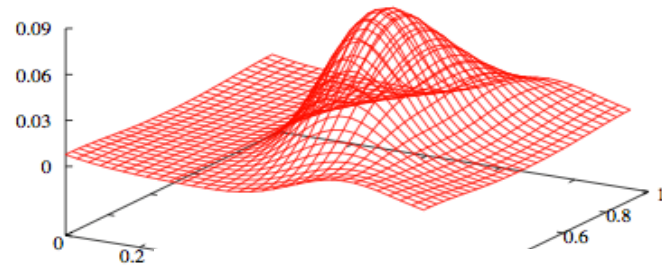
Double-null, $b < a$

Whole-space

$\Phi(t,x,y)$ at $t=2.1$



$\Phi(t,x,y)$ at $t=5$

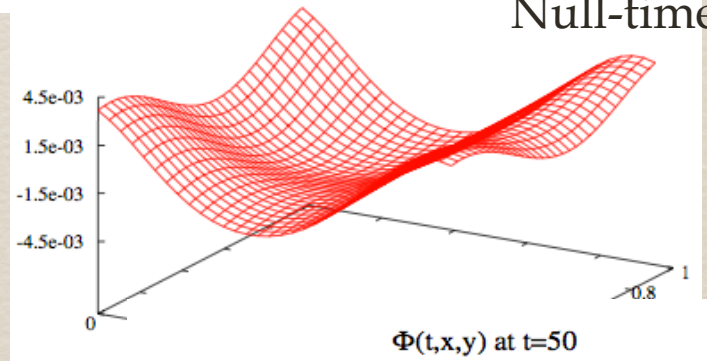
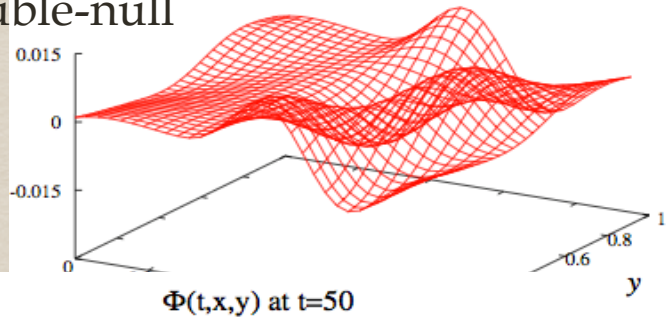


$\Phi(t,x,y)$ at $t=10$

$\Phi(t,x,y)$ at $t=10$

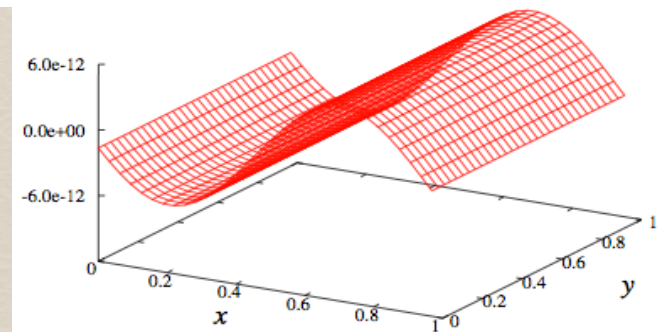
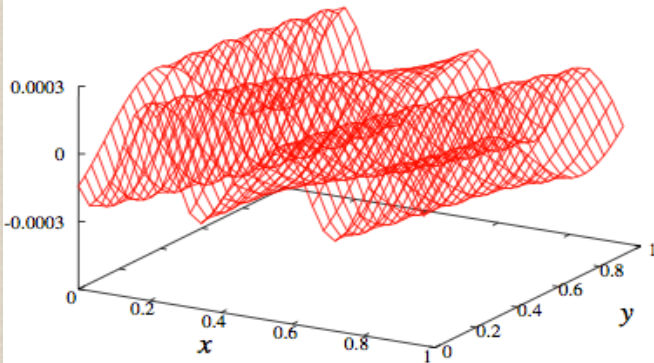
Null-timelike

Double-null



$\Phi(t,x,y)$ at $t=50$

$\Phi(t,x,y)$ at $t=50$



Source

$$ds^2 = -(e^{2\beta}W - r^{-2}h_{AB}W^A W^B)dt^2 \quad * \text{ Bondi-Sachs metric}$$

$$-2e^{2\beta}dtdr - 2h_{AB}W^B dt dx^A + r^2 h_{AB} dx^A dx^B$$

$$(2\partial_t \partial_r - W \partial_r^2)(r\Phi) = r(\partial_r W) \partial_r \Phi - r^{-1} \partial_r (W^A D_A \Phi)$$

$$+ r^{-1} D_A (e^{2\beta} D_A \Phi) - r^{-1} D_A (W^A \partial_r \Phi) + S$$

* Evolution Equation

* Compactification

$$r \rightarrow x = r/(r+R_E)$$

$$\partial_t (\partial_x \Phi + a\Phi) = \partial_x ((1-x)^2 \partial_x \Phi)$$

$$+ \partial_y^2 \Phi - 2b \partial_y \Phi + S$$

$$x=1 \Rightarrow \partial_t (\partial_x \Phi + a\Phi) = \partial_y^2 \Phi - 2b \partial_y \Phi$$

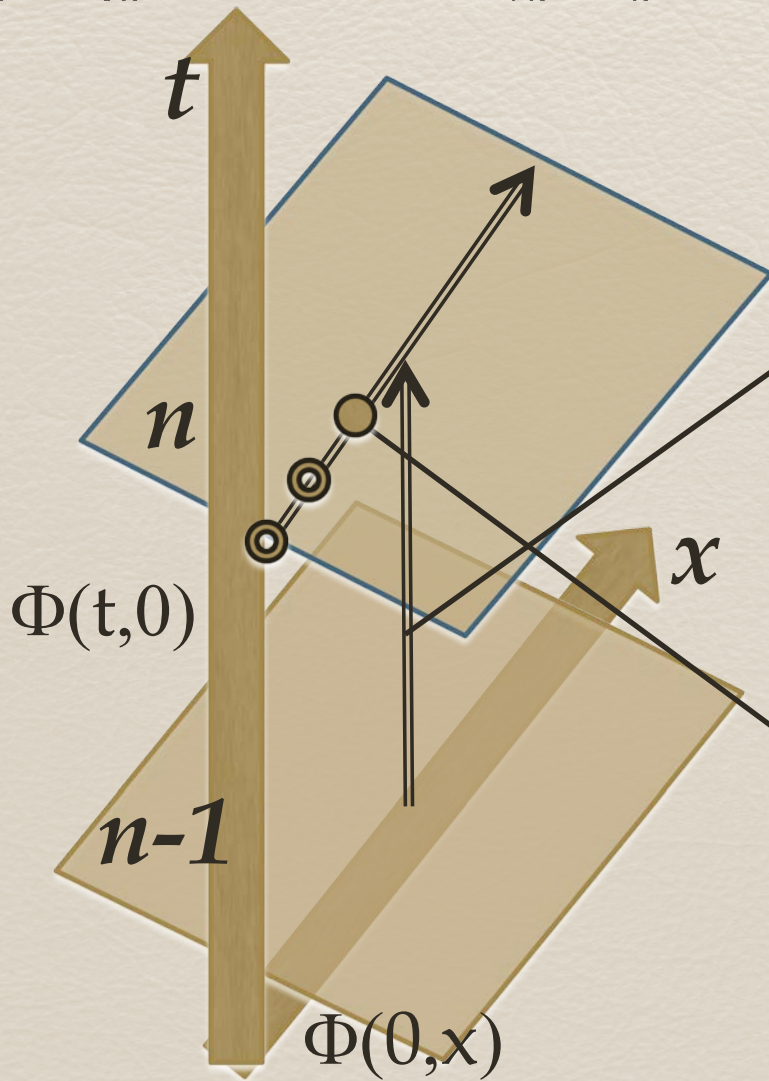
$$x=0 \Rightarrow \partial_t (\partial_x \Phi + a\Phi) = (\partial_x^2 + \partial_y^2) \Phi$$

$$-2(\partial_x + b\partial_y) \Phi$$

* Simplified model wave equation, in the null-timelike half-plane case (strip problem)

Aim

$$\partial_t(D_{0x}\Phi + a\Phi) = D_{+x}D_{-x}\Phi$$



The only stable 2-levels, marching algorithm

1. Time update

$$\Upsilon_1^n = f(\Phi_0^{n-1}, \Phi_1^{n-1}, \Phi_2^{n-1}),$$

$$\Upsilon_{j_1}^n = f(\Phi_{j_1-1}^{n-1}, \Phi_{j_1}^{n-1}, \Phi_{j_1+1}^{n-1}),$$

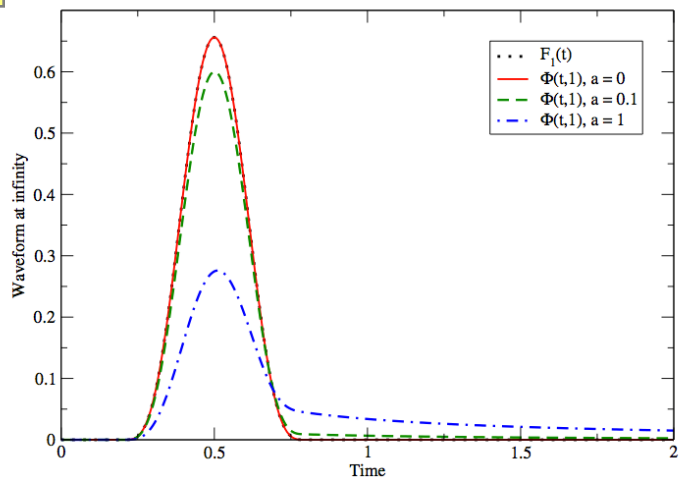
$$\Upsilon_N^n = f(\Phi_{N-1}^{n-1}, \Phi_N^{n-1}, 0)$$

2. Step-up to infinity

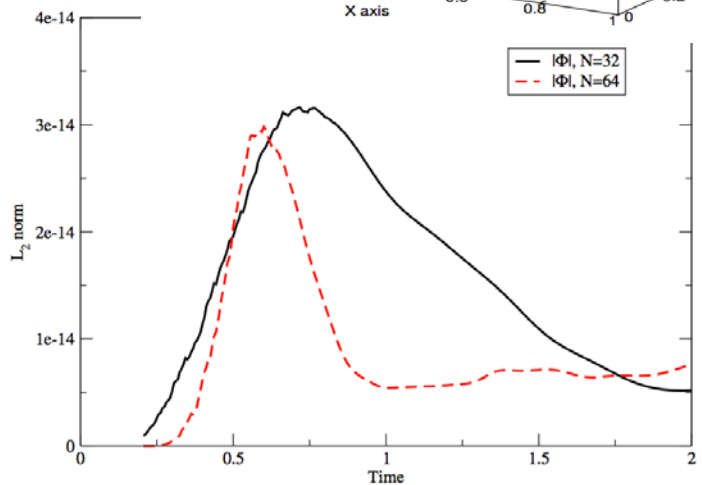
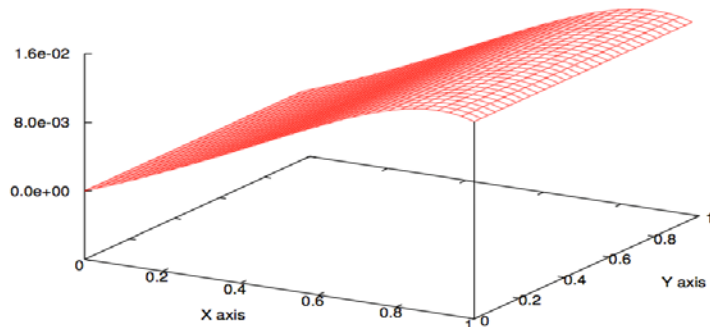
$$\Phi_1^n = f(\Upsilon_1^n, \Phi_0),$$

$$\Phi_{j_1}^n = f(\Upsilon_{j_1}^n, \Phi_{j_1-1}^n, \Phi_{j_1-2}^n)$$

Algorithm



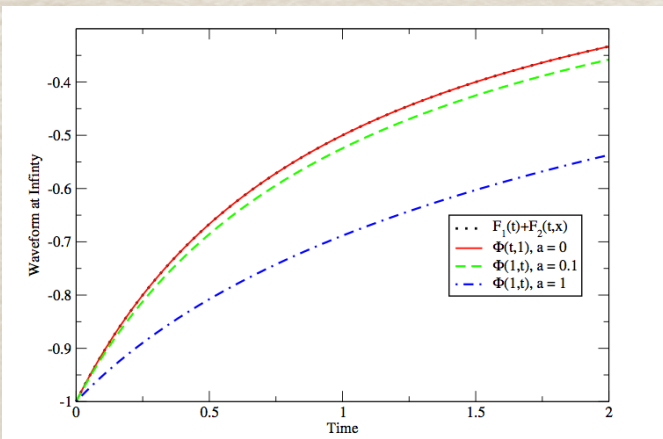
$\Phi(t,x,y)$ at $t=2$ for $a=1$



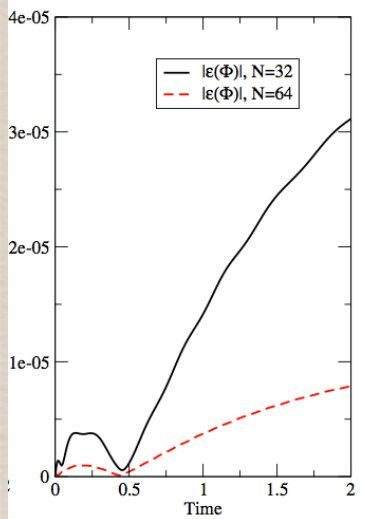
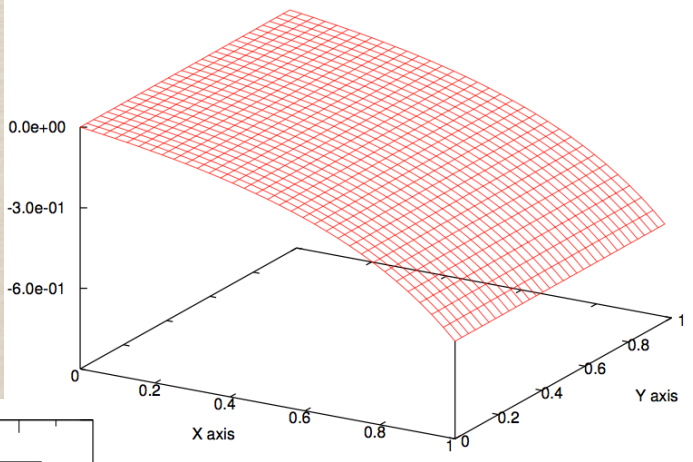
- * Purely outgoing wave
 $\Phi(0, x) = 0, \Phi(t, 0) = F_1(t)$

$$F_1(t) = A(t - t_1)^4 (t - t_2)^4$$
- * Damping due to a-term
- * Horizontal “tail” near null infinity boundary
- * Time dependence resolved exact, errors \sim machine precision

Boundary



$\Phi(t,x,y)$ at $t=2$ for $a=1$



* Purely incoming wave

$$\Phi(0, x) = -x, \quad \Phi(t, 0) = 0$$

$$F_1(t) = -\frac{1}{1+t}, \quad F_2(t) = \frac{1-x}{1+(1-x)t}$$

* Damping due to a-term

* Smooth reflection at $x=0$

* Small 2nd order errors, from discretization on x

Initial Data

$$D_{+j_2} D_{-j_2} \Phi_{j_1, j_2}^{n-1} - 2b D_{0j_2} \Phi_{j_1, j_2}^{n-1} \Rightarrow \hat{\Phi}_{j_1, j_2}^{n-1} \hat{s}_{j_1, f_2},$$

$$\varepsilon (D_{+j_2} D_{-j_2})^2 \Phi_{j_1, j_2}^{n-1} \Rightarrow \hat{\Phi}_{j_1, j_2}^{n-1} \hat{d}_{j_1, f_2},$$

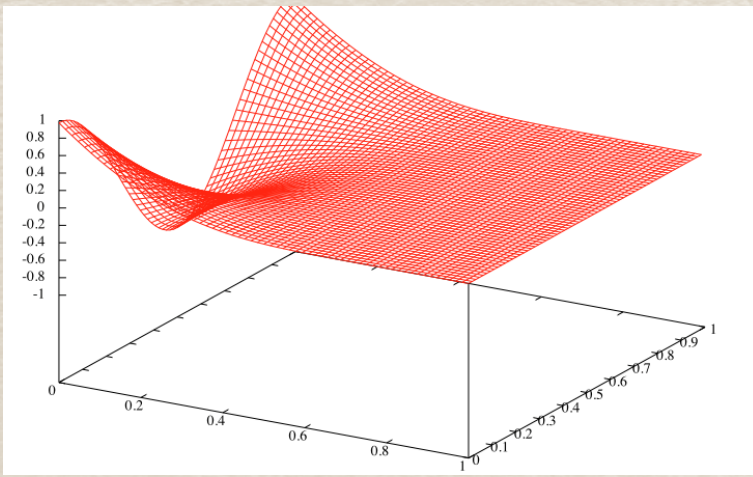
$$D_{+j_2} D_{-j_2} \Phi_{j_1, j_2}^n - 2b D_{0j_2} \Phi_{j_1, j_2}^n \Rightarrow \hat{\Phi}_{j_1, j_2}^n \hat{s}_{j_1, f_2}$$

$$\varepsilon (D_{+j_2} D_{-j_2})^2 \Phi_{j_1, j_2}^n \Rightarrow \hat{\Phi}_{j_1, j_2}^n \hat{d}_{j_1, f_2}$$

1. Initial data Fourier transformed in y
2. Add y-derivative and y-dissipation
3. Evolve in Fourier space to next level
4. Step in x direction
5. Add y-derivative and y-dissipation

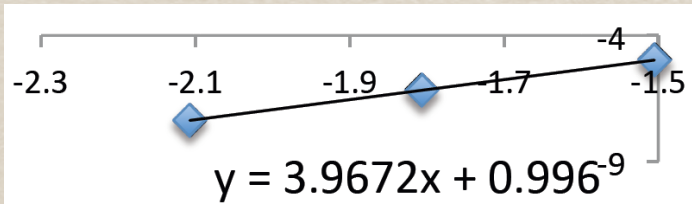
a > 0 keeps the problem well posed, for but b ≠ 0 we expect an instability, so dissipation is required

The Y-terms



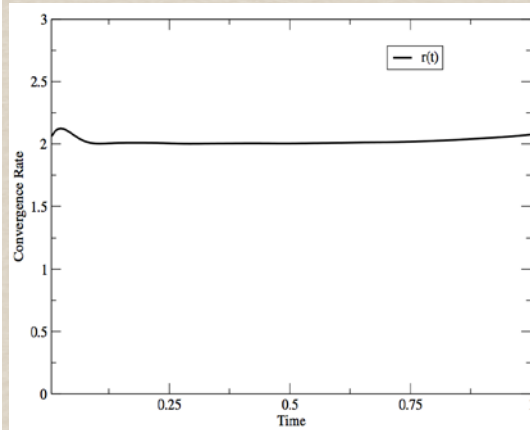
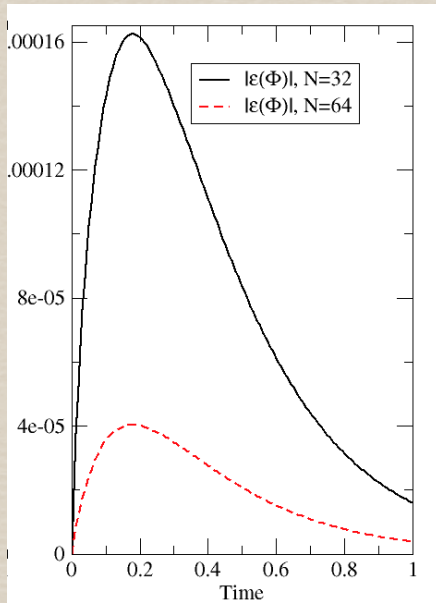
★ Exact solution

$$\Phi = e^{st} e^{sx/(1-x)} \cos(\omega y), \quad s = -\frac{\omega^2}{a}$$



★ Local 4th-order convergence

★ Global 2nd-order convergence



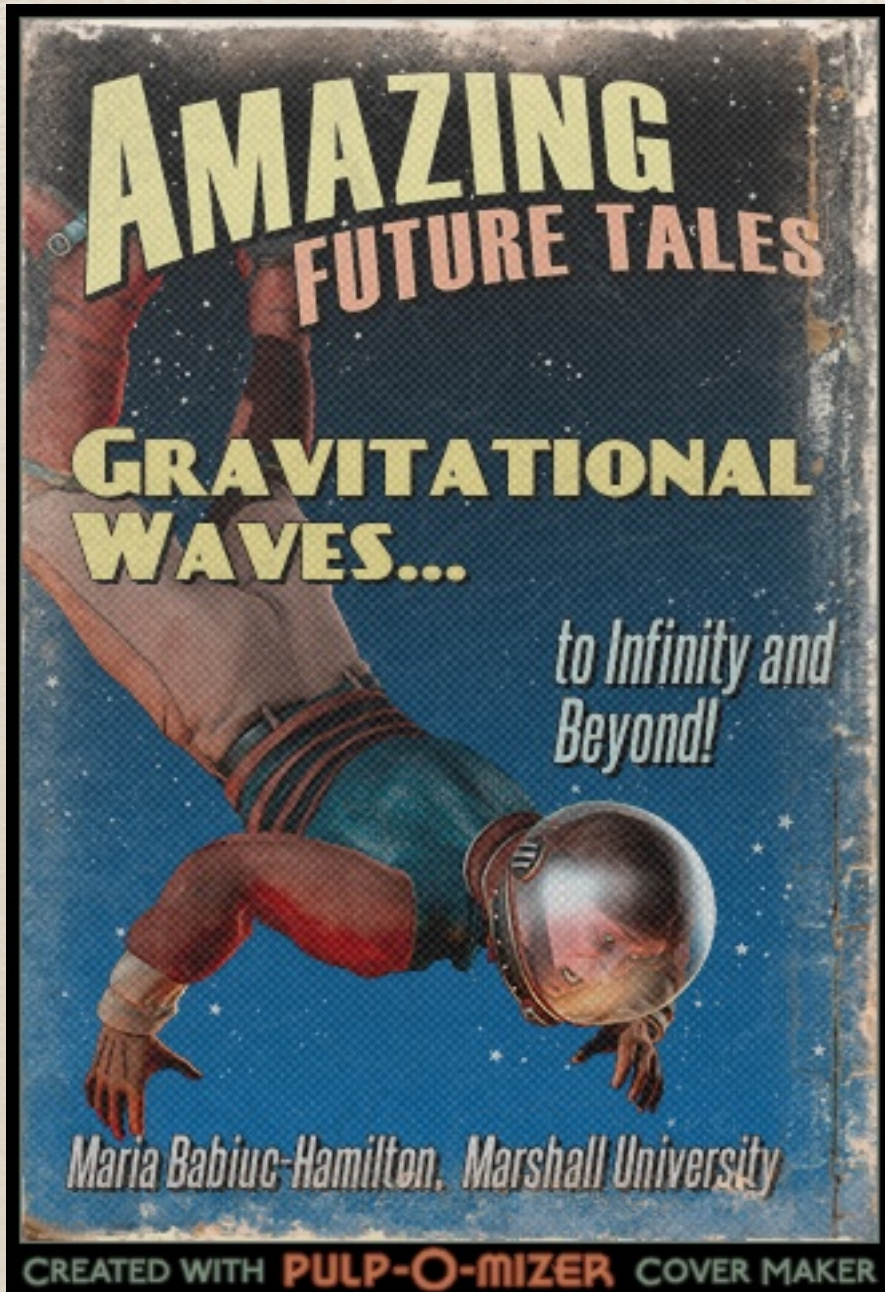
Convergence



$a=-1, b=0$
Ill-Posed

- * The proof for well-posedness was done analyzing the behavior of continuous and discrete exponential Fourier-Laplace modes as solutions of the wave equations in characteristic coordinates
- * Well-posedness is ensured for correct choice of parameters (a,b)
- * For the whole space, $a > 0$ renders the runs stable and convergent
 - * the double-null system still has exponentially growing modes
 - * the null-timelike system has no growing modes for $|b| < a$
- * For the half-plane, due to the term $(1-x)$, there is a range for a
 - * at $x=0$, weak strong condition for well-posedness for $a > -2$
 - * at $x=1$, strong strong condition for well-posedness for $a > 0$
- * For $b^2 < a(a+2)$ there are no growing modes. There is an inherent instability for $b \neq 0$, and angular dissipation is necessary.

Conclusions



- * This proof of well-posedness of the corresponding problem for the quasilinear wave equation is a first step toward treating the fully nonlinear gravitational case.

* Jeff Winicour, Pittsburgh Univ.

* NSF Grant PHY -0969709

Thank You