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Adding Light to the Gravitational Waves on the Null Cone

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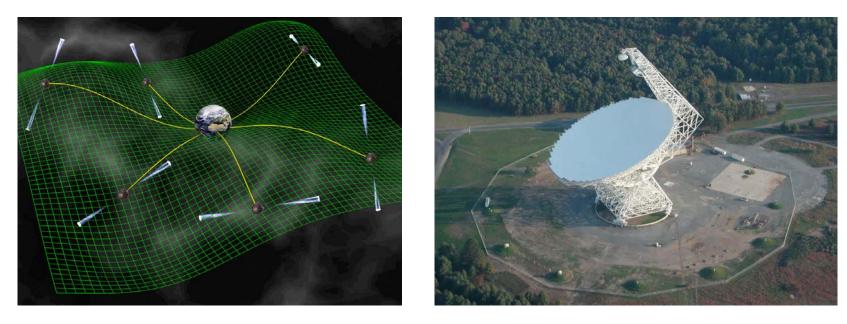
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ADDING LIGHT TO THE GRAVITATIONAL WAVES ON THE NULL CONE

Maria Babiuc Marshall University, WV APS Physics APRIL Meeting 2014 April 5-8, Savannah, Georgia

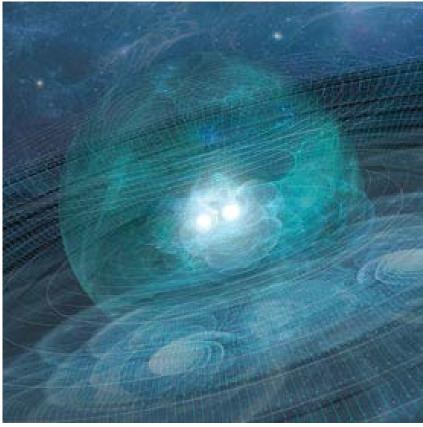
PTA and the search for GW



- The Pulsar Timing Array (PTA) will be able to detect not only stochastic GW, but also resolve individual super massive binaries through their EM emission.
- SMBH binaries reside in recent galaxies mergers, of massive, nearby host galaxies, with possible AGN.

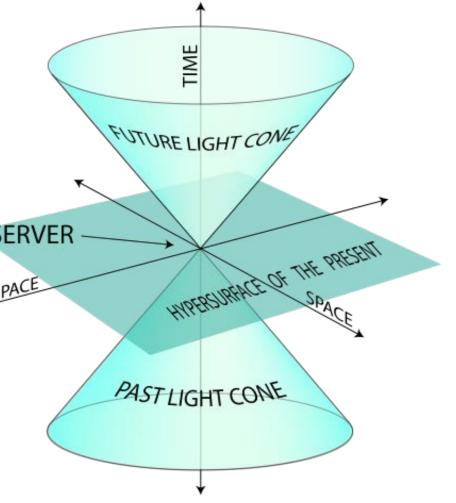
We Can See Gravitational Waves

- If the merger of two SMBHs produce a light signal.
- The accretion disks around merging BHs, emit EM radiation that includes the GW signature of the binary.
- Periodic flares on a timescale correlated with the GW frequency, or even the birth of a bright variable quasar, could indicate the presence of SMBH binaries.



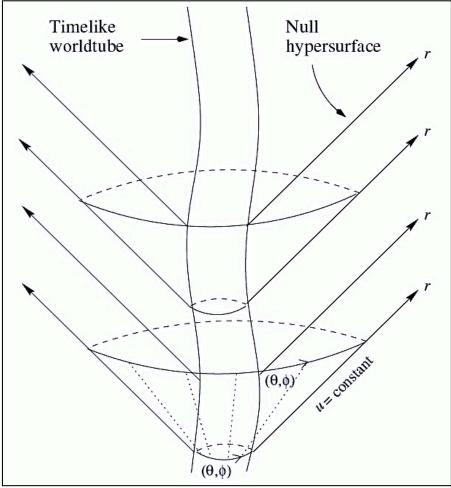
How Radiation Travels

- Both gravitational and electromagnetic radiation travel along light rays, which are *principal null directions* in space-time.
- They are *characteristic* OBSERVER *surfaces* both of Einstein's vacuum field equations and SPACE of Maxwell's equations.
- In characteristic coordinates Einstein equations split in evolution and hypersurface.



The Characteristic Code

- Einstein equations $G_{\mu\nu}=0$ are propagated radially and in time along the outgoing light rays, in **Bondi-Sachs** coordinates, by a marching integration algorithm.
- The gravitational waveforms are computed at positive null infinity on inertial Bondi coordinates in terms of the Bondi news and Weyl scalar.



Enlighten the Gravity

• Cover the whole space-time with the Sachs metric:

$$ds^{2} = \left(-e^{2\beta}\frac{V}{r} + r^{2}U^{2}e^{2\psi} + r^{2}W^{2}e^{-2\psi}\right)du^{2} - 2e^{2\beta}dudr - 2r^{2}We^{2\psi}dud\theta$$

 $-2r^2Y\sin\theta e^{-2\psi}dud\varphi+r^2(e^{2\psi}d\theta^2+e^{-2\psi}\sin^2\theta d\varphi^2)$

• Consider an EM field described by the Faraday tensor:

 $F = F_{01}du \wedge dr + F_{02}du \wedge d\theta + F_{03}du \wedge d\varphi + F_{12}dr \wedge d\theta + F_{13}dr \wedge d\varphi + F_{23}d\theta \wedge d\varphi$

• Write down the field equations to be solved:

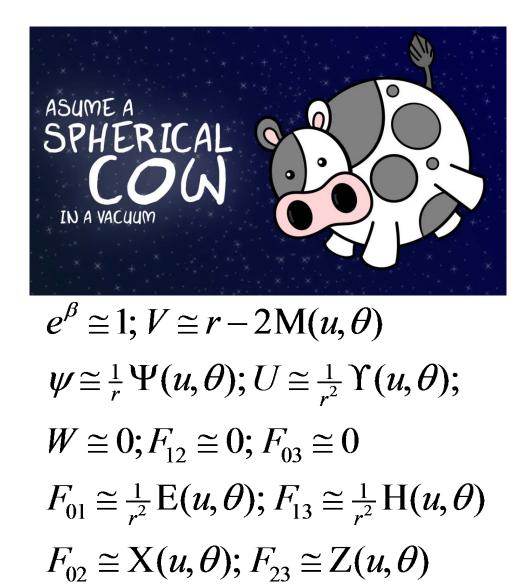
$$G_{\alpha\beta} = 8\pi T_{\alpha\beta}; \quad D_{[\delta}F_{\alpha\beta]} = 0; \quad D^{\alpha}F_{\alpha\beta} = 0$$

$$G_{\alpha\beta} = R_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}R; \quad T_{\alpha\beta} = \frac{1}{4\pi} \left(F_{\alpha}^{\delta}F_{\beta\delta} - \frac{1}{4}g_{\alpha\beta}F_{\delta\sigma}F^{\delta\sigma}\right)$$

$$D^{\alpha}G_{\alpha\beta} = 0; \quad D^{\alpha}T_{\alpha\beta} = 0; \quad D^{\alpha}W_{\alpha\beta\delta\sigma} = \frac{1}{2} \left(D_{\delta}R_{\beta\sigma} - D_{\sigma}R_{\beta\delta}\right)$$

Make Reasonable Assumptions

- Causality: no inflow radiation from future.
- Vacuum: the Ricci scalar curvature is zero.
- Axial symmetry will not restrict the generality.
- Linearity: The parallax, and luminosity distance r along the rays agree.
- Euclidean topology: the space-time is flat at a distance reasonably far away from the source.



Book Keeping

- The linearized Sachs metric with axial symmetry is:
- $ds^{2} = \left(-1 + \frac{2M}{r} + \Upsilon^{2}e^{\frac{2}{r}\Psi}\right)du^{2} 2dudr 2r^{2}\Upsilon e^{\frac{2}{r}\Psi}dud\theta + r^{2}(e^{\frac{2}{r}\Psi}d\theta^{2} + e^{-\frac{2}{r}\Psi}\sin^{2}\theta d\phi^{2})$ The Faraday tensor for the electromagnetic field is:

$$F = F_{01}du \wedge dr + F_{02}du \wedge d\theta + F_{13}dr \wedge d\varphi + F_{23}d\theta \wedge d\varphi$$

• The field equations split into twelve main equations:

$$R_{01} - 8\pi T_{01} = 0; \quad R_{12} - 8\pi T_{12} = 0; \quad R_{13} - 8\pi T_{13} = 0$$

$$R_{22} - 8\pi T_{22} = 0; \quad R_{23} - 8\pi T_{22} = 0; \quad R_{33} - 8\pi T_{33} = 0$$

$$D_{[2}F_{01]} = 0; \quad D_{[3}F_{01]} = 0; \quad D_{[3}F_{12]} = 0$$

$$D^{\alpha}F_{\alpha 0} = 0; \quad D^{\alpha}F_{\alpha 2} = 0; \quad D^{\alpha}F_{\alpha 3} = 0$$

• The unknown variables are: M, Y, Ψ , E, H, Z, X

The Questions are:

- 1. Can we still split all these equations into evolution and hypersurface equations for the main variables?
- 2. Can we decouple the gravitational and the EM fields?
- Ideally: with initial data for the gravitational field Ψ_0 , and for the electromagnetic field (E_0 , H_0), we calculate (Y_0 , M_0 , Z_0 , X_0) and integrate to (Ψ_1 , E_1 , H_1).
- A hierarchical integration uses hypersurface equations for (Y, M, Z, X) and evolution equations for (Ψ, E, H).
- Hierarchical equations:
 - 1. R_{12} -8 πT_{12} =0=>Y₀ function of Z₀
 - 2. $R_{22}=8\pi T_{12}=0$ =>M₀ function of Z₀
 - 3. $R_{33}=8\pi T_{33}=0=>Z_0=>Y_0, M_0$
 - 4. $J_2 = 0 = X_0$
- Evolution equations: *R*₀₁=8π*T*₀₁ ->Ψ₁; *J*₀=0 ->E₁; *J*₃=0 ->H₁

Numerical Algorithm

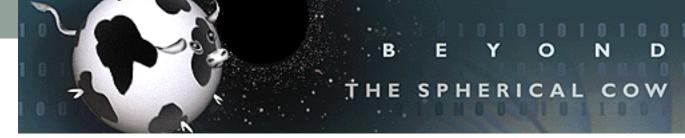
- The hierarchical equations contain only angular derivatives and will be integrated in Fourier space.
- The evolution equations can be integrated with a generic 4th order (Runge-Kutta) integrator, because the dependence on the extraction radius *r* is explicit.
- Construct the Weyl tensor, taking into account the EM

$$W_{\alpha\beta\delta\sigma} = R_{\alpha\beta\delta\sigma} - \frac{1}{2} \left(g_{\alpha\delta} E_{\beta\sigma} + g_{\beta\sigma} E_{\alpha\delta} - g_{\alpha\sigma} E_{\beta\delta} - g_{\beta\delta} E_{\alpha\sigma} \right)$$

$$E_{\alpha\beta} = R_{\alpha\beta} - 8\pi T_{\alpha\beta}$$

 Both the Weyl and Faraday fields will have to be projected onto a quasi-normal tetrad to extract the scalars encoding the measurable values of the fields.

What's up?



- Future extensions of the code:
 - Drop the axial symmetry
 - Extend the power series expansion in r
 - Add the radial dependence
 - Compactify the *r* coordinate
 - Change the integration algorithm
 - Carry the evolution to null infinity
 - Do the asymptotic expansion to a inertial coordinate system
 - Calculate the Weyl scalar and Poynting flux
 - Include null neutrino dust
 - Include universe expansion

A New Characteristic Code

- Numerical relativity groups could develop their own homemade characteristic extraction modules.
- Electromagnetic counterparts of gravitational waves can point to gravitational sources.
- Other interesting phenomenal
 - gravitational memory effect
 - formation of trapped surfaces
 - the problem of horizon

The Gravitational Wave Spectrum

